

# Final Exam (Spring 2017)

## PHYS 510: Classical Mechanics

Date: 2017 May 11

1. (20 points.) A catenary is the curve that an idealized hanging chain assumes under its own weight when supported only at its ends in a uniform gravitational field. It is the curve  $y(x)$  that minimizes the potential energy of the hanging chain. Let us assume the two end points of the chain are at the same height. A catenary is given by

$$y = a \cosh \frac{x}{a}, \quad (1)$$

where the parameter  $a$ , an integration constant, characterizes the catenary.

- (a) Determine the perimeter length  $P$  of the hanging chain using

$$P = \int_{-x_0}^{x_0} ds. \quad (2)$$

- (b) Find the relation between the parameter  $a$ , the perimeter length  $P$  of the chain, and the height  $y_0$  in Figure 1.

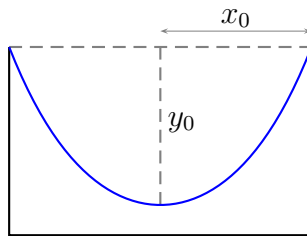


Figure 1: Problem 1.

2. (40 points.) Kepler problem is described by the potential energy

$$U(r) = -\frac{\alpha}{r}, \quad (3)$$

and the corresponding Lagrangian

$$L(\mathbf{r}, \mathbf{v}) = \frac{1}{2}\mu v^2 + \frac{\alpha}{r}. \quad (4)$$

For the case when the total energy  $E$  is negative,

$$-\frac{\alpha}{2r_0} < E < 0, \quad r_0 = \frac{L_z^2}{\mu\alpha}, \quad (5)$$

where  $L_z$  is the angular momentum, the motion is described by an ellipse,

$$r(\phi) = \frac{r_0}{1 + e \cos(\phi - \phi_0)}, \quad e = \sqrt{1 + \frac{E}{(\alpha/2r_0)}}. \quad (6)$$

Perihelion is the point in the orbit of a planet at which it is closest to the Sun. The precession of the perihelion is suitably defined in terms of the displacement  $\Delta\phi$  of the perihelion during one revolution,

$$\Delta\phi = -2\pi + 2 \int_{r_{\min}}^{r_{\max}} d\phi, \quad (7)$$

where  $r_{\min}$  is the perihelion when the planet is closest to Sun and  $r_{\max}$  is the aphelion when the planet is farthest from Sun. Show that the precession of perihelion is zero for the Kepler problem.

When a small correction

$$\delta U(r) = \frac{\beta}{r^2} \quad (8)$$

is added to the potential energy the precession of the perihelion is no longer zero. Expanding in powers of  $\delta U$  determine the precession of the perihelion to the leading order for this correction.

3. **(20 points.)** The Poincaré formula for the addition of (parallel) velocities is,  $c = 1$ ,

$$v = \frac{v_a + v_b}{1 + v_a v_b}, \quad (9)$$

where  $v_a$  and  $v_b$  are velocities and  $c$  is speed of light in vacuum. Assuming that the Poincaré formula holds for all speeds, subluminal ( $-1 < v_i < 1$ ), superluminal ( $|v_i| > 1$ ), and speed of light, analyse the addition of a subluminal speed and a superluminal speed? That is, is the ‘sum’ subluminal or superluminal. Is the answer unique?

4. **(20 points.)** The path of a relativistic particle moving along a straight line with constant (proper) acceleration  $\alpha$  is described by the equation of a hyperbola

$$x^2 - c^2 t^2 = x_0^2, \quad x_0 = \frac{c^2}{\alpha}. \quad (10)$$

This is the motion of a particle ‘dropped’ from  $x = x_0$  at  $t = 0$  in region of constant (proper) acceleration. Will a photon dispatched to ‘chase’ this particle at  $t = 0$  from  $x = 0$  ever catch up with it? If yes, when and where does it catch up?